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Consolidated Picture of MagLIF implosions & Current State of our Understanding

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Our current physical picture of the stagnated fuel and liner is based on a series of x-ray and neutron diagnostics

14 12 Axial Position [mm] 10 8 6 -1.0 -0.5 0.0 0.5 1.0 Radial Pos. [mm]

X-ray image of

MagLIF plasma

Stagnating plasma $<Te> = 3 \pm 1 \text{ keV}$ $<Ti> = 2.5 \pm 0.5 \text{ keV}$ $T(r) = T_0[1 - (r/R)^2]$ $t_{burn} = 1.5 \pm 0.5 \text{ ns}$ $f_{mix} \sim 1 \%$ $R = 50 \pm 20 \ \mu\text{m}$ $\rho_D = 0.3 \pm 0.1 \ \text{g/cm}^3$ $P(z) = 1 \pm 0.2 \ \text{Gbar}$ $\rho R \sim 1.5 \ \text{mg/cm}^2$ $BR \sim 0.4 \ \text{MG} \cdot \text{cm}$

Confining liner $\rho r_{liner} = 1 \text{ g/cm}^2$

- Stagnation (DD) column is weakly helical and contiguous in z along most of the length of target (although significant temperature variations can exist)
- Stagnated DD is highly magnetized (BR ~ 0.4 MG-cm) with 25-55% magnetic flux conservation
- ~40-50% (sim) of the fuel mass is conserved in the implosion, ~10-20% inferred from *hot-spot* emission volume
- Relatively slow (70 km/s) and stable (Δdr(z)<<r) Be liner. Outer liner material stagnates on fuel 5 ns after peak neutron production of the DD fuel
- High CR, could be as high as 45!





Why do our simulations, in general, predict higher performance than observed?

























We are collecting data on all phases of MagLIF implosions, on multiple facilities



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Our current focus is on better understanding of fuel preheating and mix



